

The Effect of Atmospheric Resistance, Magnetic Force and Oblateness of the Earth on the Motion and Stability of two satellites connected by an extensible cable in circular orbit of the Centre of Mass.

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Abstract: In a linear motion of a system of two satellites connected by extensible cable, one stable equilibrium point exists when perturbative forces like Solar radiation pressure, shadow of the earth, oblateness of the earth, air resistance and earth's magnetic force act simultaneously. We have obtained one stable point of equilibrium in case of perturbative forces like atmospheric resistance, Magnetic Force and oblateness of the earth acting together on the motion of a system of two satellites connected by extensible cable in the central gravitational field of earth in case of circular orbit of the centre of mass. We have used Liapunov's theorem on stability to examine the stability of the equilibrium point.

Keywords: Stability, Equilibrium Point, circular orbit, Liapunov Theorem, Satellites, perturbative forces.

1. Introduction

The present paper is concerned with the stability of the equilibrium point of the centre of mass of a system of two satellites connected by a light, flexible and extensible cable under the influence of Atmospheric Resistance, Magnetic Force and oblateness of the earth in case of circular orbit). Beletsky, V.V is the pioneer worker in this field. This paper is an attempt towards the generalization of works done by him.

2. Equation of motion of the system

The equations of motion of one of the two satellites moving along Keplerian elliptic orbit under the influence Atmospheric Resistance, Magnetic Force and oblateness of the earth in Nechvill's coordinates system may be obtained by exploiting Lagrange's equations of motion of first kind in the form:

$$\begin{aligned}
 x'' - 2y' + 3x\rho + \frac{4A}{\rho} + f\rho\rho' + \frac{B}{\rho} \cos i &= -\bar{\lambda}_\alpha \left[1 - \frac{l_0}{\rho_r} \right] x\rho^4 \\
 y'' + 2x' + f\rho^2 - \frac{A}{\rho} y + \frac{B\rho'}{\rho^2} \cos i &= -\bar{\lambda}_\alpha \rho^4 \left[1 - \frac{l_0}{\rho_r} \right] y \\
 \text{and } z'' + z - \frac{A}{\rho} z &= -\bar{\lambda}_\alpha \rho^4 \left[1 - \frac{l_0}{\rho_r} \right] z - \frac{B}{\rho} \left[\frac{\rho'}{\rho} \cos(v+w) + \frac{f}{\mu_E} (3p^3 \rho^2 - \mu_E) \sin(v+w) \right] \sin i \dots\dots\dots(2.1)
 \end{aligned}$$

Equation of constraint is given by

$$x^2 + y^2 + z^2 \leq \frac{l_0^2}{\rho^2} \dots\dots\dots(2.2)$$

Where,

$$r = \sqrt{x^2 + y^2 + z^2}, A = -\frac{3k_2}{\rho^2}, B = \left(\frac{Q_1}{m_1} - \frac{Q_2}{m_2} \right) \frac{\mu_E}{\sqrt{\mu p}}, f = \frac{ap^3}{\sqrt{\mu p}} \quad \& \quad \bar{\lambda}_\alpha = \frac{p^3}{\mu} \lambda_\alpha = \frac{p^3}{\mu} \left(\frac{m_1 + m_2}{m_1 m_2 l_0} \right) \lambda$$

For the circular orbit of the centre of mass of the system, we must have

$$e = 0, \rho = \frac{1}{1 + e \cos v} = 1 \quad \& \quad \rho' = 0$$

On putting $\rho = 1, \rho' = 0$ & $i = 0$ in the equation of motion (1), we get the new system of equation in the form

$$\begin{aligned}
 x'' - 2y' - (3 - 4A)x + B &= -\bar{\lambda}_\alpha \left[1 - \frac{l_0}{r} \right] x \\
 y'' + 2x' + f - Ay &= -\bar{\lambda}_\alpha \left[1 - \frac{l_0}{r} \right] y \\
 \& \quad z'' + z - Az = -\bar{\lambda}_\alpha \left[1 - \frac{l_0}{r} \right] z \quad \dots\dots\dots(2.3)
 \end{aligned}$$

The condition of constraints takes the form

$$x^2 + y^2 + z^2 \leq l_0^2 \quad \dots\dots\dots (2.4)$$

Multiplying 1st, 2nd and 3rd equations of (3) by $2x'$, $2y'$ and $2z'$ respectively and adding we get

$$\left. \begin{aligned}
 2x'x'' + 2y'y'' + 2z'z'' - (3 - 4A)2xx' + 2Bx' + 2fy' - 2Ayy' + 2(1 - A)zz' \\
 + \bar{\lambda}_\alpha (2xx' + 2yy' + 2zz') + \bar{\lambda}_\alpha l_0 \frac{(2xx' + 2yy' + 2zz')}{\sqrt{x^2 + y^2 + z^2}} = 0
 \end{aligned} \right\} \dots\dots\dots(2.5)$$

Integrating (2.5) we get Jacobian integral in the form,

$$x^2 + y^2 + z^2 - (3 - 4A)x^2 + 2Bx + 2fy - Ay^2 + (1 - A)z^2 + \bar{\lambda}_\alpha (x^2 + y^2 + z^2) + \bar{\lambda}_\alpha l_0 (x^2 + y^2 + z^2)^{\frac{1}{2}} = h \quad \dots\dots\dots(2.6)$$

Where h is the constant of integration

3. Equilibrium solution of the problem

The equilibrium position of the system is given by the constant values of the co-ordinates in rotating frame of references.

Let,

$$x = x_0, y = y_0, z = z_0 \quad \dots\dots\dots(3.1)$$

$$\therefore x' = x_0' = 0 = x''$$

$$y' = y_0' = 0 = y'' \quad \dots\dots\dots (3.2)$$

$$z' = z_0' = 0 = z''$$

Where x_0 , y_0 and z_0 are constants give equilibrium position of the system .

Using (3.1) and (3.2) in (2.3) we get

$$\left. \begin{aligned}
 -(3 - 4A)x_0 + B &= -\bar{\lambda}_\alpha \left(1 - \frac{l_0}{r_0} \right) x_0 \\
 f - Ay_0 &= -\bar{\lambda}_\alpha \left(1 - \frac{l_0}{r_0} \right) y_0 \\
 (1 - A)z_0 &= -\bar{\lambda}_\alpha \left(1 - \frac{l_0}{r_0} \right) z_0
 \end{aligned} \right\} \dots\dots\dots(3.3)$$

where, $r_0 = \sqrt{x_0^2 + y_0^2 + z_0^2}$

Now, we shall discuss the particular solutions of the system of equations (3.3) as follows:

From first and second equations of (3.3) it follows that magnetic force parameter B and atmospheric resistance parameter f vanish if $x_0 = 0$ and $y_0 = 0$. Hence $x_0 \neq 0$ and $y_0 \neq 0$. If we put $z_0 = 0$ in the third equations of (3.3) then it is not possible to find the values of $x_0 = 0$ and $y_0 = 0$ from the first two equations of (3.3)

Therefore, possible equilibrium position for the system of equations (2.3) may be taken as (a,b,c)

Putting $x_0 = a$ and $y_0 = b$ and $z_0 = c$ in equation (3.3) we get

$$\left. \begin{aligned} -(3-4A)a+B &= -\bar{\lambda}_\alpha \left(1-\frac{l_0}{r_0}\right)a \\ f-Ab &= \bar{\lambda}_\alpha \left(1-\frac{l_0}{r_0}\right)b \\ (1-A)c &= -\bar{\lambda}_\alpha \left(1-\frac{l_0}{r_0}\right)c \end{aligned} \right\} \dots\dots\dots(3.4)$$

Where, $r_0 = \sqrt{a^2 + b^2 + c^2}$

From the last equation of (3.4) we have

$$1-A = -\bar{\lambda}_\alpha \left(1-\frac{l_0}{r_0}\right) \text{ as } (c \neq 0) \dots\dots\dots(3.5)$$

Using (3.5) in the first equation and 2nd equation of (3.4) we get

$$x_0 = a = \frac{B}{4-5A} \text{ and } y_0 = b = f \dots\dots\dots (3.6)$$

From (3.5) we get

$$\frac{\bar{\lambda}_\alpha l_0}{r_0} = \bar{\lambda}_\alpha + 1 - A$$

$$\therefore r_0 = \frac{\bar{\lambda}_\alpha l_0}{\bar{\lambda}_\alpha + 1 - A}$$

$$i.e. \sqrt{a^2 + b^2 + c^2} = \frac{\bar{\lambda}_\alpha l_0}{\bar{\lambda}_\alpha + 1 - A}$$

$$\therefore c = \pm \sqrt{\left(\frac{\bar{\lambda}_\alpha l_0}{\bar{\lambda}_\alpha + 1 - A}\right)^2 - \left(\frac{B}{4-5B}\right)^2 - f^2}$$

Hence the equilibrium position in case of circular orbit of the centre of mass of the system is given by

$$(a, b, c) = \left\{ \frac{B}{4-5B}, f, \sqrt{\left(\frac{\bar{\lambda}_\alpha l_0}{\bar{\lambda}_\alpha + 1 - A}\right)^2 - \left(\frac{B}{4-5A-f^2}\right)^2} \right\} \dots\dots\dots(3.7)$$

4. Stability of the system

Now, we proceed to test the stability of the equilibrium position given by (3.7).

Let η_1, η_2 and η_3 denote the small variations in the given equilibrium position. Then we have

$$\left. \begin{aligned} x &= a + \eta_1, y = b + \eta_2, \text{ and } z = c + \eta_3 \\ x' &= \eta_1', y' = \eta_2', \text{ and } z' = \eta_3' \\ x'' &= \eta_1'', y'' = \eta_2'', \text{ and } z'' = \eta_3'' \end{aligned} \right\} \dots\dots\dots(4.1)$$

Using (4.1) in (2.3) we get the variational equations of motion for the system in the form

$$\left. \begin{aligned} \eta_1'' - 2\eta_2' - (3-4A)(a + \eta_1) + B &= -\bar{\lambda}_\alpha \left[1-\frac{l_0}{r}\right](a + \eta_2) \\ \eta_2'' + 2\eta_1' + f - A(b + \eta_2) &= -\bar{\lambda}_\alpha \left[1-\frac{l_0}{r}\right](b + \eta_2) \\ \eta_3'' + (1-A)(c + \eta_3) &= -\bar{\lambda}_\alpha \left[1-\frac{l_0}{r}\right](c + \eta_3) \end{aligned} \right\} \dots\dots\dots(4.2)$$

$$\text{Where, } r^2 = (a + \eta_1)^2 + (b + \eta_2)^2 + (c + \eta_3)^2 \dots\dots\dots(4.3)$$

As the original equations of motion given by (2.3) admits Jacobi's integral, so the variational equations of motion given by (4.2) will also admit Jacobi's integral and can be easily obtained by multiplying the three equations of (4.2) by $2(a + \eta_1)$, $2(b + \eta_2)$ and $2(c + \eta_3)$ respectively and adding them together and integrating in the form.

$$\eta_1^2 + \eta_2^2 + \eta_3^2 - (3 - 4A)(a + \eta_1)^2 + 2f(a + \eta_2) - A(b + \eta_2)^2 + (1 - A)(c + \eta_3)^2 + 2B(a + \eta_1) + \bar{\lambda}_\alpha \left[(a + \eta_1)^2 + (b + \eta_2)^2 + (c + \eta_3)^2 - 2\bar{\lambda}_\alpha l_0 r \left[1 + \left\{ 2(a\eta_1 + b\eta_2 + c\eta_3) + \frac{\eta_1^2 + \eta_2^2 + \eta_3^2}{r^2} \right\}^{\frac{1}{2}} \right] \right] = h \tag{4.4}$$

Where h is the constant of integration. Equation (4.4) gives us the Jacobi's integral for the variational equations of motion of the system at the equilibrium position (a,b,c)

To examine the stability in the sense of Liapunov, we take Jacobian integral as Liapunov's function

$V(\eta_1, \eta_2, \eta_3, \eta_1, \eta_2, \eta_3)$ and is obtained by expanding the terms in (4.4) as

$$V(\eta_1, \eta_2, \eta_3, \eta_1, \eta_2, \eta_3) = \eta_1^2 + \eta_2^2 + \eta_3^2 + \eta_1^2 \left[\bar{\lambda}_\alpha + 4A - B - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 a^2}{r^3} \right] + \eta_2^2 \left[\bar{\lambda}_\alpha - A - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 b^2}{r^3} \right] + \eta_3^2 \left[\bar{\lambda}_\alpha + 1 - A - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 c^2}{r^3} \right] + \eta_1 \left[-2a(3 - 4A) + 2B + 2\bar{\lambda}_\alpha a - \frac{2l_0 \bar{\lambda}_\alpha a}{r} \right] + \eta_2 \left[2f - 2Ab + 2\bar{\lambda}_\alpha b - \frac{2\bar{\lambda}_\alpha l_0 b}{r} \right] + \eta_3 \left[2(1 - A)c + 2\bar{\lambda}_\alpha c - \frac{2\bar{\lambda}_\alpha l_0 c}{r} \right] + 2\eta_1 \eta_2 \left[\frac{\bar{\lambda}_\alpha l_0 ab}{r^3} \right] + 2\eta_2 \eta_3 \left[\frac{\bar{\lambda}_\alpha l_0 bc}{r^3} \right] + 2\eta_3 \eta_1 \left[\frac{\bar{\lambda}_\alpha l_0 ca}{r^3} \right] + o(3) \tag{4.5}$$

Where O(3) stands for the third and higher order terms in the small quantities $\eta_1, \eta_2,$ and η_3 .

By Liapunov's theorem on stability it follows that the only criterion for given equilibrium position (a, b, c) to be stable is that V defined by (4.5) must be positive definite and for this the following conditions must be satisfied:

$$\begin{aligned} \text{(i)} \quad & -2(3 - 4A)a + 2B + 2\bar{\lambda}_\alpha a - \frac{2\bar{\lambda}_\alpha l_0 a}{r} = 0 \\ \text{(ii)} \quad & 2f - 2Ab + 2\bar{\lambda}_\alpha b - \frac{2\bar{\lambda}_\alpha l_0 b}{r} = 0 \\ \text{(iii)} \quad & 2(1 - A)c + 2\bar{\lambda}_\alpha c - \frac{2\bar{\lambda}_\alpha l_0 c}{r} = 0 \end{aligned} \quad \text{(iv)} \quad \begin{vmatrix} A_1 & \frac{\bar{\lambda}_\alpha l_0 ab}{r^3} & \frac{\bar{\lambda}_\alpha l_0 ca}{r^3} \\ \frac{\bar{\lambda}_\alpha l_0 ab}{r^3} & A_2 & \frac{\bar{\lambda}_\alpha l_0 bc}{r^3} \\ \frac{\bar{\lambda}_\alpha l_0 ac}{r^3} & \frac{\bar{\lambda}_\alpha l_0 bc}{r^3} & A_3 \end{vmatrix} > 0 \tag{4.6}$$

Where,

$$\left. \begin{aligned} A_1 &= \bar{\lambda}_\alpha + 4A - 3 - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 a^2}{r^3} \\ A_2 &= \bar{\lambda}_\alpha - A - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 b^2}{r^3} \\ A_3 &= \bar{\lambda}_\alpha + 1 - A - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 c^2}{r^3} \end{aligned} \right\} \tag{4.7}$$

$$(v) \left| \begin{array}{cc} A_1 & \frac{\bar{\lambda}_\alpha l_0 ab}{r^3} \\ \frac{\bar{\lambda}_\alpha l_0 ab}{r^3} & A_2 \end{array} \right| > 0 \quad \dots\dots\dots(4.8)$$

$$(vi) A_1 > 0 \quad \dots\dots\dots(4.9)$$

From (4.6), (4.8) and (4.9) we conclude that $A_1 > 0$, $A_2 > 0$ and $A_3 > 0$. Hence finally the sufficient condition for stability of the equilibrium position (a, b, c) of the system can be put in the form.

$$\begin{aligned} (i) \quad & \left[-2a(3-4A) + 2B + 2\bar{\lambda}_\alpha a - \frac{2\bar{\lambda}_\alpha l_0 a}{r} \right] = 0 \\ (ii) \quad & \left[2f - 2Ab + 2\bar{\lambda}_\alpha b - \frac{2\bar{\lambda}_\alpha l_0 b}{r} \right] = 0 \\ (iii) \quad & \left[2 - (1-A)c + 2\bar{\lambda}_\alpha c + \frac{2\bar{\lambda}_\alpha l_0 c}{r} \right] = 0 \\ (iv) \quad & \left[A_1 = \bar{\lambda}_\alpha + 4A - 3 - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 a^2}{r^3} \right] > 0 \\ (v) \quad & \left[A_2 = \bar{\lambda}_\alpha - A - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 b^2}{r^3} \right] > 0 \\ (vi) \quad & \left[A_3 = \bar{\lambda}_\alpha + 1 - A - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 c^2}{r^3} \right] > 0 \end{aligned} \quad \dots\dots\dots(4.10)$$

Now let us proceed to examine the conditions mentioned above in (4.10) one by one to have a clear picture of the stability of the system at the equilibrium point (a, b, c)

Condition (I):

$$L.H.S. = 2 \left[B - (3-4A)a + \bar{\lambda}_\alpha \left(1 - \frac{l_0}{r} \right) a \right] = 2 \left[B - (3-4A)x_0 + \bar{\lambda}_\alpha \left(1 - \frac{l_0}{r} \right) x_0 \right] = 2.0 = 0 =$$

R.H.S

$$(II): L.H.S = 2 \left[f - Ab + \bar{\lambda}_\alpha \left(1 - \frac{l_0}{r} \right) b \right] = 2 \left[f - Ay_0 + \bar{\lambda}_\alpha \left(1 - \frac{l_0}{r} \right) y_0 \right] = 2.0 = 0 = R.H.S$$

[Using the second equation of (3.4)]

Condition (III)

$$L.H.S = 2 \left[(1-A)c + \bar{\lambda}_\alpha \left(1 - \frac{l_0}{r} \right) c \right] = 2 \left[(1-A)z_0 + \bar{\lambda}_\alpha \left(1 - \frac{l_0}{r} \right) z_0 \right] = 2.0 = 0 = R.H.S.$$

[Using the third equation of (3.4)]

Condition (IV)

$$A_1 = \bar{\lambda}_\alpha + 4A - 3 - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 a^2}{r^3} = \bar{\lambda}_\alpha + 4A - 3 - \frac{\bar{\lambda}_\alpha l_0 [\bar{\lambda}_\alpha + 1 - A]}{\bar{\lambda}_\alpha l_0} + \frac{\bar{\lambda}_\alpha l_0}{r^3} a^2 = 5A - 4 + \frac{\bar{\lambda}_\alpha l_0}{r} (a/r)^2$$

$$= 5(A-1) + \frac{1 + \bar{\lambda}_\alpha l_0}{r} (a/r)^2 = 5(A-1) + 1 + (\bar{\lambda}_\alpha + 1 - A)(a/r)^2 > 0$$

$$\therefore r = \frac{\bar{\lambda}_\alpha l_0}{\bar{\lambda}_\alpha + 1 - A} > 0 \Rightarrow \bar{\lambda}_\alpha + 1 - A > 0$$

Hence fourth condition is satisfied if $A-1 > 0$

Condition (V):

$$A_2 = \bar{\lambda}_\alpha - A - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 b^2}{r^3} = \bar{\lambda}_\alpha - A - \frac{\bar{\lambda}_\alpha l_0 [\lambda \alpha^2 + 1 - A]}{\bar{\lambda}_\alpha l_0} + \frac{\bar{\lambda}_\alpha l_0}{r} (b/r)^2 = -1 + \frac{\bar{\lambda}_\alpha l_0}{r} \frac{b^2}{r^2}$$

$$= (\bar{\lambda}_\alpha + 1 - A) \frac{b^2}{r^2} - 1 > 0$$

$$\text{if } \frac{r^2}{b^2} < \bar{\lambda}_\alpha + 1 - A \Rightarrow \frac{1}{b^2} < \frac{\bar{\lambda}_\alpha + 1 - A}{r^2}$$

Hence fifth condition is satisfied if $\frac{1}{f^2} < \frac{\bar{\lambda}_\alpha + 1 - A}{r^2}$

Condition (VI)

$$A_3 = \bar{\lambda}_\alpha + 1 - A - \frac{\bar{\lambda}_\alpha l_0}{r} + \frac{\bar{\lambda}_\alpha l_0 c^2}{r^3} = \bar{\lambda}_\alpha + 1 - A - \frac{\bar{\lambda}_\alpha l_0 [\bar{\lambda}_\alpha + 1 - A]}{\bar{\lambda}_\alpha l_0} + \frac{\bar{\lambda}_\alpha l_0}{r} \frac{c^2}{r^2} = (\bar{\lambda}_\alpha + 1 - A) \frac{c^2}{r^2} > 0$$

$$\left[\because (\bar{\lambda}_\alpha + 1 - A > 0, \text{ as } r > 0) \right]$$

Hence sixth condition is also satisfied. Thus we see that the six conditions of (4.10) for the stability of the equilibrium position (a, b, c) are identically satisfied if

$$A > 1 \text{ and } -\frac{\sqrt{\bar{\lambda}_\alpha + 1 - A}}{r} < \frac{1}{b} < \frac{\sqrt{\bar{\lambda}_\alpha + 1 - A}}{r}$$

Hence we conclude that the equilibrium is stable at (a, b, c) in the sense of Liapunov.

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